## **Event processing in X-IFU detector onboard Athena**

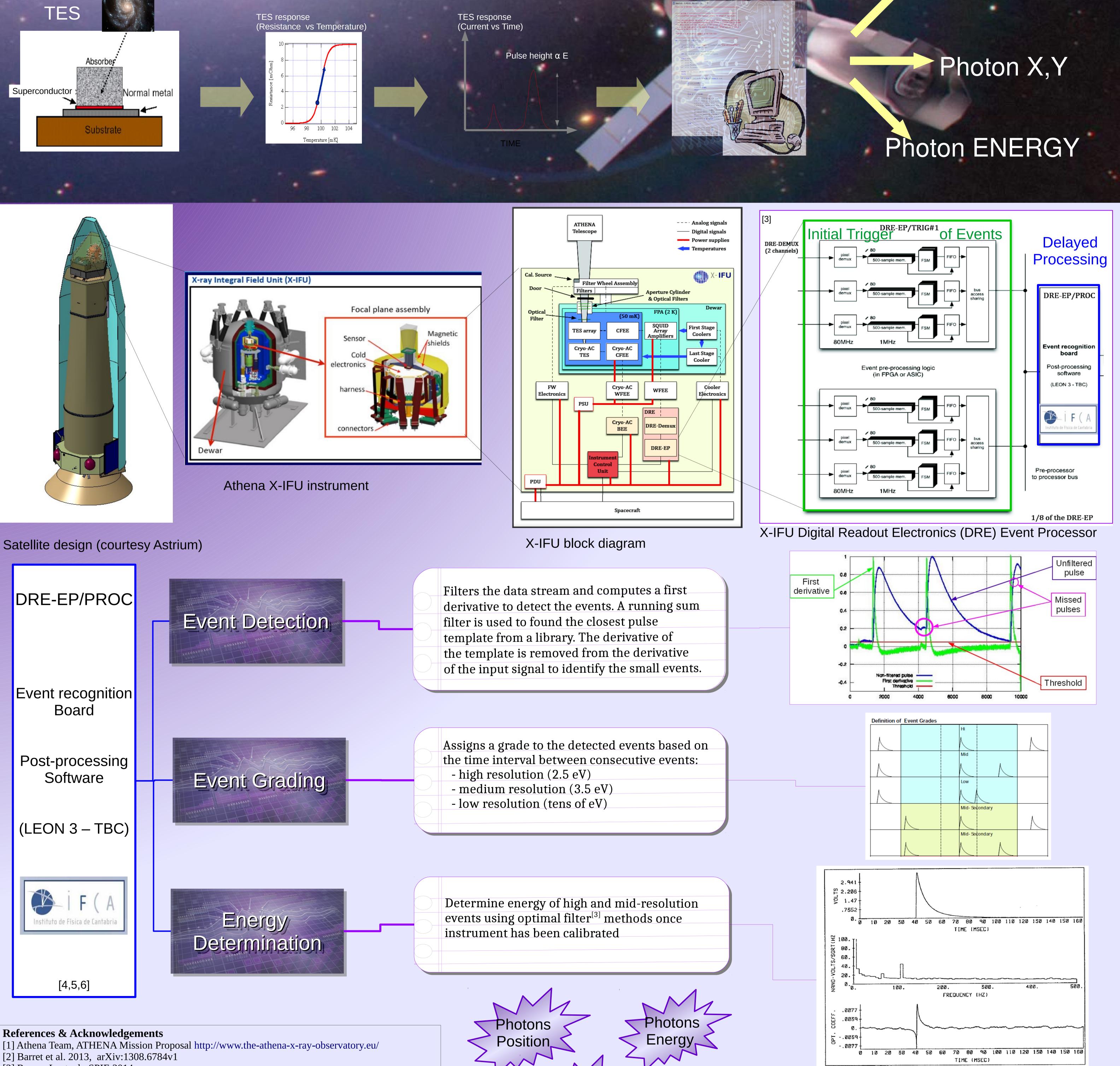
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New Astronomy: The current and forthcoming research lines in X-ray astronomy (black holes, accretion physics, hot cosmic plasmas, large bound structures) will require unprecedented spectral resolution with imaging capabilities. New Mission: In June 2014, the X-ray observatory ATHENA<sup>[1]</sup> was selected by ESA to implement the science theme "The Hot and Energetic Universe" as the second large-class mission. New Detectors: The X-IFU<sup>[2]</sup> (X-ray Integral Field Unit) onboard ATHENA is a calorimeter based on Transition Edge Sensor (TES) technologies, able to provide high spectral resolution (2.5 eV @ 6 keV) New processing: These new detectors require a different approach for the event detection: they must detect the electrical pulses that are the response to an abrupt change in resistance in the device, caused by the absorption of an X-ray photon.

Let's see how to process this...!











[1] Athena Team, ATHENA Mission Proposal http://www.the-athena-x-ray-observatory.eu/ [2] Barret et al. 2013, arXiv:1308.6784v1 [3] Ravera L. et al. SPIE 2014 [4] Ceballos M.T. et al. 2011, ASP Conference Series, Vol 442, 335 [5] Ceballos M.T. et al. 2012, ASP Conference Series, Vol 461, 777 [6] Ceballos M.T. et al. 2013, ASP Conference Series, Vol. 475, 25 [7] Szyimkowiak A.E. et al. 1993, JLTP, Vol. 93, Nos 3/4 This work has been funded by the Spanish Ministry of Science and Innovation (MICINN) under projects ESP2006-13608-C02-01, AYA2009-08059, AYA2010-21490-C02-01, AYA2012-39767-C02-01

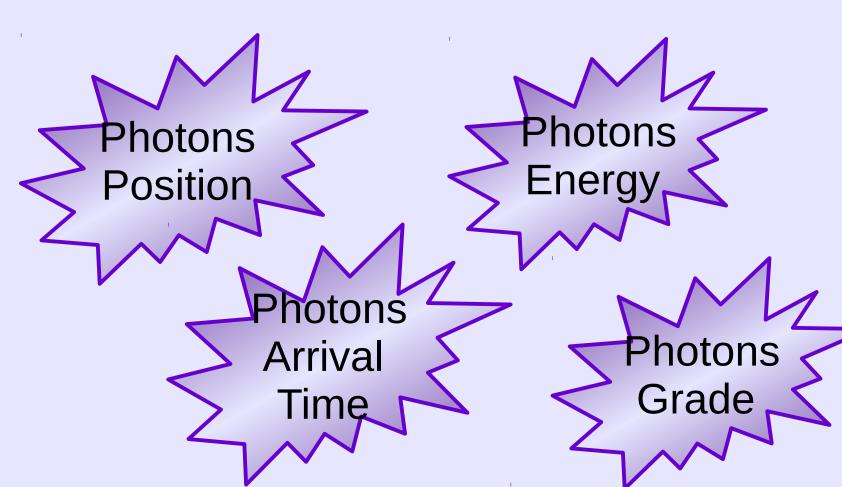


Fig. 1 Plots showing (top to bottom) average pulse shape, power spectrum of the noise, and the optimal filtering template. [7]